Development and Performance of X-ray Astronomical SOI pixel sensor

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Posters Ryu et al. "Trigger-Driven Readout" NP3.M-90 Nakashima et al. "Analog to Digital Converter" NP3.M-92





Nakashima-kun

Papers

Ryu et al. IEEE NSS 2010, Conf. Record (2010) Ryu et al. IEEE TNS 58, 2528 (2011) Nakashima et al. TIPP 2011 (2011) Submitted

20111024_XRPIX_SOIPIX_IEEE-NSS11_v11.key

The standard Imaging Spectrometer of X-ray CCD

Suzaku「すざく」 XIS

- Fano limited spectroscopy with the readout noise ~3e- (rms).
- Wide and fine imaging with the sensor size of ~20-30mm pixel size of ~30µm[□]

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- Non X-ray background above I0keV is too high to study faint sources.
- The time resolution is too poor (~ sec) to make fast timing obervation of variable sources.

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"XRPIX" = SOI pixel sensor for future X-ray astronomical satellites



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XRPIX1-CZ (X-Ray PIXel detector - CZochralski)







- 17 keV and 8 keV X-rays have different attenuation lengths.
- Measure the depletion thickness by observing the ratio between the counting rates of the two energies X-rays.
- The data follow the expectation well.

Nakashima et al., 2011, submitted



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XRPIX1-FZ (7kΩcm): Dark Current from Rough Backside Surface



XRPIX1-FZ (7kΩcm): Dark Current from Rough Backside Surface



XRPIX1-FZ (7kΩcm): Dark Current from Rough Backside Surface





- The thickness of the depletion layer of XRPIXI-FZ reaches
 ~250µm at 30V and stops its growth there.
- The 250 μ m is nearly equal to the hi- ρ Si thickness (260 μ m).
- Full depletion is achieved at VBB=30V.

Nakashima et al., 2011, submitted

<u>XRPIXI-CZ : X-ray Spectra in frame mode</u>



Ryu et al., 2011

XRPIXI-CZ : Readout Noise

- See if the readout noise of 100e- (rms) is explained by the sum of circuit noises or not.
- Measure the noise of individual circuit element through several DC voltage input points.

The sum of these noise are almost consistent with the observed readout noise of 100e.



$\frac{\mathsf{XRPIXI} \to \mathsf{XRPIXIb}}{\mathsf{XRPIXIb}}$

Purpose

Improvement of Spectroscopic performance

Block I

- Increase C_CDS from 100fF to 400fF to reduce the reset noise generated at the CDS circuit.
- Block 2, 3
 - Reduce the area of BPW to 45% to obtain higher gain.
 - BPW (buried p-well) suppressing back gate effect dominates the capacitance at the sensor node.

Show results from block 2 today





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$\underline{\mathsf{XRPIXI}}_{\mathsf{CZ}} \rightarrow \mathsf{Ib}_{\mathsf{CZ}} : \underline{\mathsf{X}}_{\mathsf{ray}} \text{ Spectra in frame mode}$



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- Reduce noises other than the reset noise by introducing LPF. high_v(100 samples average) - low_v(100 samples average) \rightarrow LPF with τ =100µs





- When X-ray is detected, the device outputs trigger signal.
- Address of the triggered pixel is output according to clock from FPGA.
- ADC reads out the analog signal of only the triggered pixel.

XRPIX-ADCI



800 Data from Odd modulator 600 $EIN = 72 \mu V rms$ Counts 400 200 Offset Q 10 12 16 18 20 22 14 Digital output (ADU)

Fig. 6. Histogram of digital output in the case of GND input.

- $\Delta\Sigma$ type (over sampling type)
- SOI version of the ASIC developed for X-ray CCD camera onboard ASTRO-H (next Japanese X-ray satellite).
- It is working.
- Making performance test now.



See Nakashima's poster (NP3.M-92) in detail.

XRPIX2 : New Device





Designed by

- Large Size, Large Format
- 60µm□ : Single pixel can cover the whole charge cloud to reduce charge sharing effect.
- Make the capacitance at the readout node smaller (Area of BPW = 1/4 of XRPIX1).
- Make further increase of C_CDS to reduce the reset noise.
- Just submitted last week.



Thank you.





