Performance of SMILE-2+ Electron-Tracking Compton Camera

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MeV gamma-ray Astronomy



Requirements for the next-generation observation are

- Wide-band detection
- Large Field of View
- High quality image

Problem and Answer of difficulties



Well-defined Point Spread Function (PSF) and Low BG are necessary for MeV gamma-ray telescope

Electron-Tracking Compton Camera (ETCC)



Gaseous Time Projection Chamber

Tracker: 3-D track and energy of electrons

2D imaging (x,y) + Drift Time (z) =>3D

Pixel Scintillator Arrays

Absorber: absorption position and

energy of scattered gamma-ray



Reconstruct Compton scattering event by event completely

- Back ground rejection
- Two-Dimensional PSF

deposited energy

- \bullet Particle identify with dE/dX = -
- track length
- \blacklozenge Compton Kinematical test with α

➤Without heavy VETO

>Large FOV (~ 3 str)



Sub-MeV gamma-ray Imaging Loaded-on-balloon Experiment

SMILE-I @ Sanriku (Sep. 1st 2006)

> We obtained diffuse cosmic/atmospheric gamma-ray spectra

Success in rejection with dE/dX A. Takada+. ApJ,2011

SMILE-2: Only ground test T. Tanimori+. ApJ,2015

SMILE-2+ 1-day flight @ Alice Springs (Apr. 7th 2018)
 Certification of imaging spectroscopy in MeV Astrophysics
 observation of bright objects (Crab & Galactic Center)
 Requirement Effective area a few cm²
 (detect 5σ) PSF (50% included) ~ 10 deg @ 662 keV

SMILE-3

> Scientific observation loaded on a long duration balloon



All sky survey with an ETCC loaded on a satellite ~ sub-mCrab sensitivity

To improve effective area of SMILE-2+



(1)

3





used as a molecular sieve





High Energy Electron Event





Next generation require BG rejection & Well defined PSF
We are developing ETCC to meet the requirements
We developed SMILE-2+ as hard as possible
to make shift with launching balloon
Now we are doing best to analyze the data of ground test and flight data



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