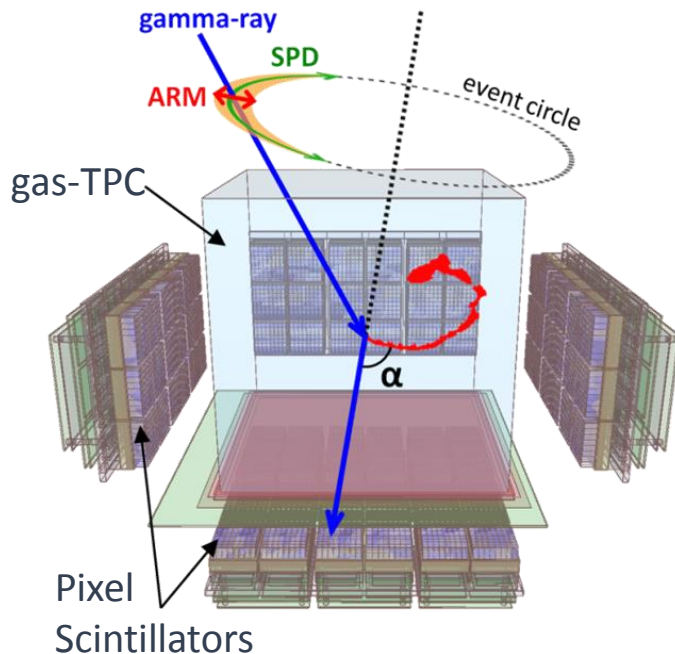


# SMILE24 : 電子飛跡検出型コンプトンカメラが持つ 偏光撮像性能の評価試験

古村翔太郎

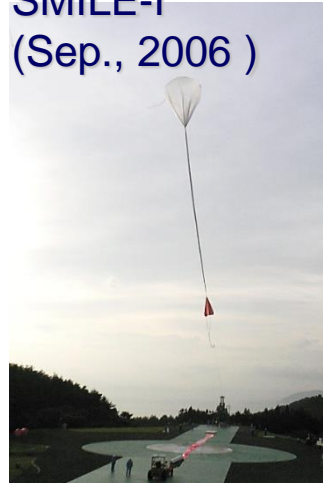
谷森達, 窪秀利, 高田淳史, Parker Joseph, 水村好貴, 水本哲矢, 園田真也, 友野大, 中村輝石,  
松岡佳大, 中村祥吾, 岸本哲郎, 小田真, 竹村泰斗, 宮本奨平, 中増勇真, 吉川慶 (京都大学)  
身内賢太郎(神戸大学), 黒澤俊介(東北大学), 澤野達哉(金沢大学)



**Sub-MeV gamma-ray Imaging**

**Loaded-on-balloon Experiment**

**SMILE-I**  
(Sep., 2006)



**Now planning**

SMILE-II (1 day flight)

Crab Nebula, Cygnus X-1

SMILE-III (1 month flight)

- more faint sources
- polarization measurements

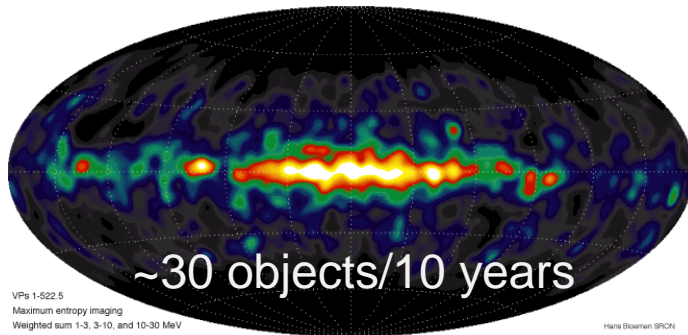
# Contents

---

1. MeVガンマ線天文学の現状・課題
  2. ETCCのコンセプト
  3. Current and near-future ETCC
  4. 偏光検出器としての応用
  5. 偏光ビーム試験@SPring-8 2015年1月
- ◆ Summary

# Problems in MeV gamma-ray astronomy

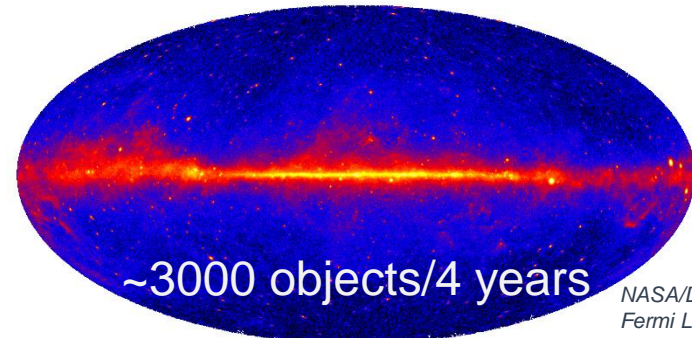
COMPTEL 1-30 MeV map



~30 objects/10 years

V. Schönfelder+ (A&AS, 2000)

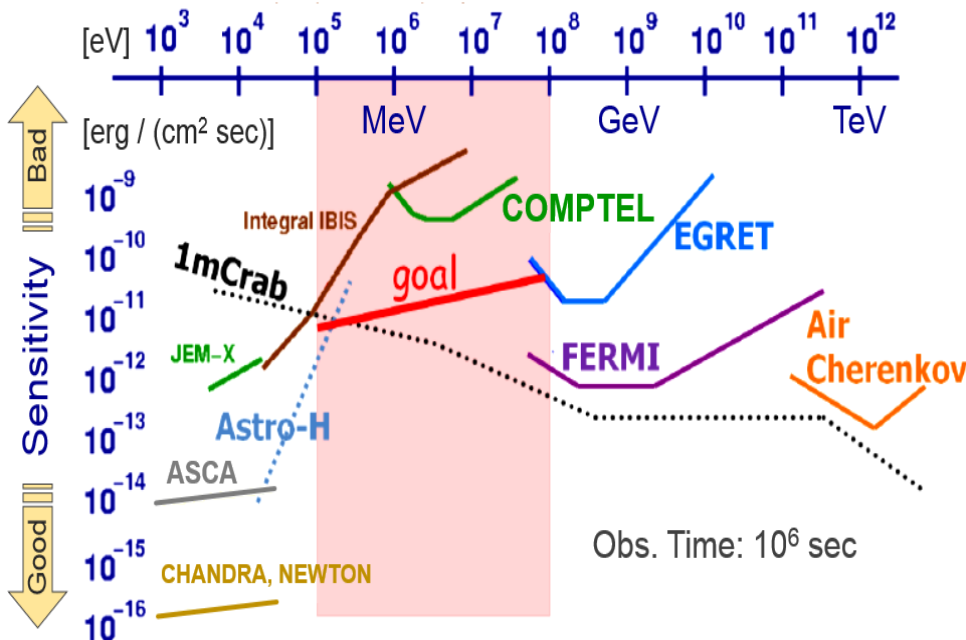
Fermi > 1 GeV map



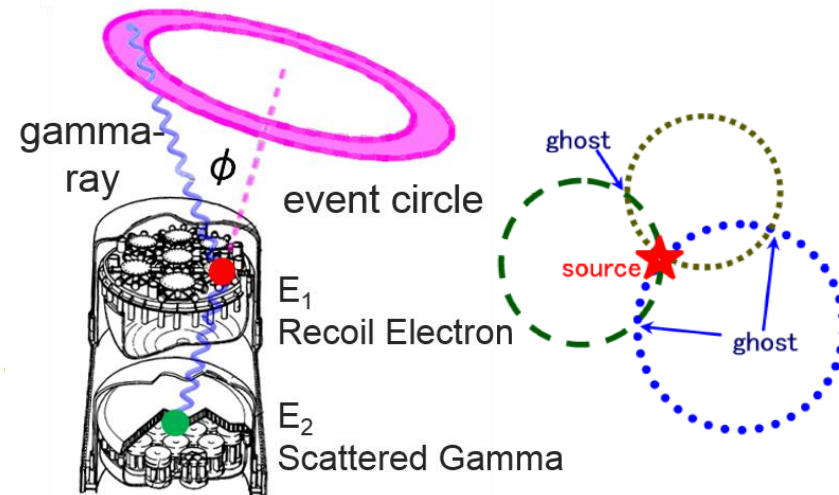
~3000 objects/4 years

NASA/DOE/  
Fermi LAT Collaboration

F. Acero+ (ApJS, 2015)

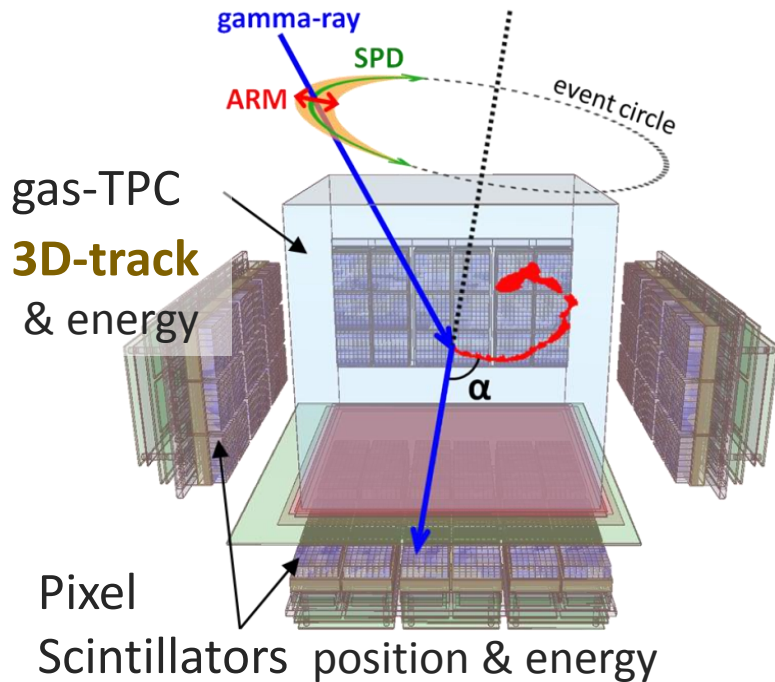


Unexplored frontier!



- ◆ Wide-spread PSF
- ◆ Huge backgrounds in space

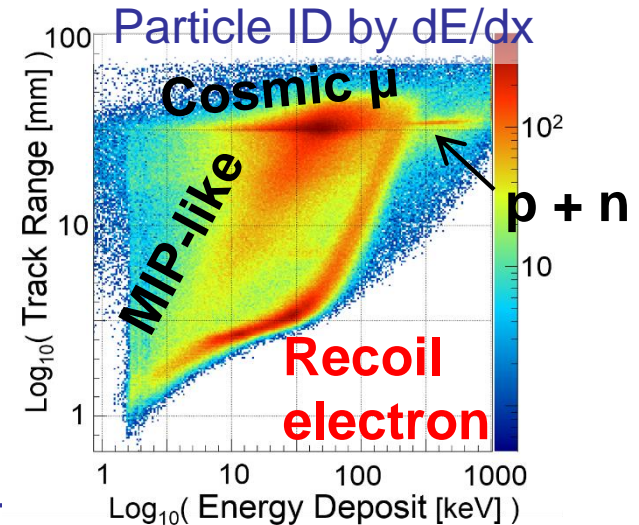
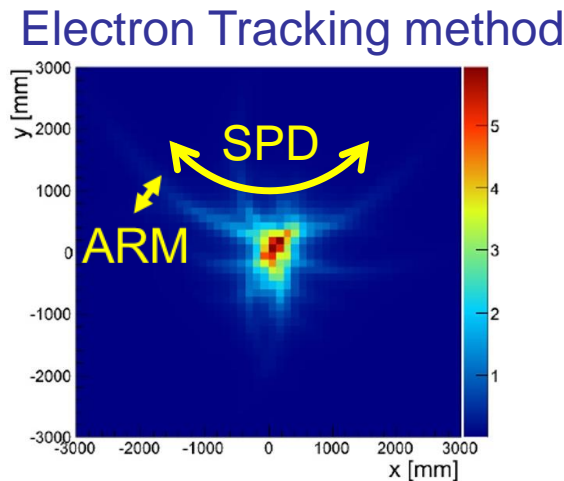
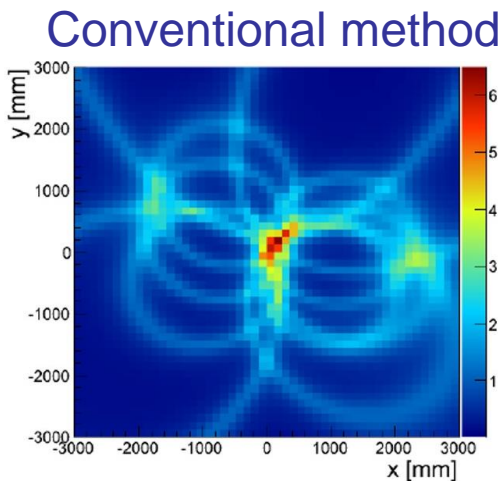
# Electron-Tracking Compton Camera (ETCC)



By measuring electron tracks, ETCC overcome the problems !

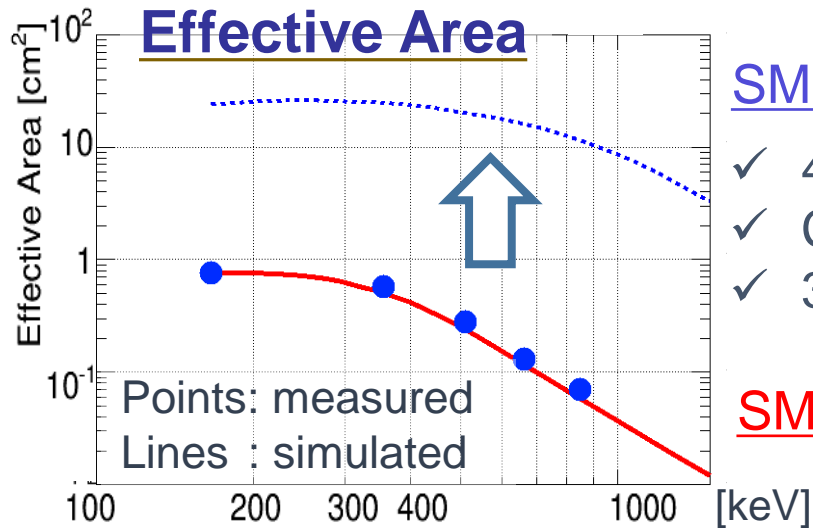
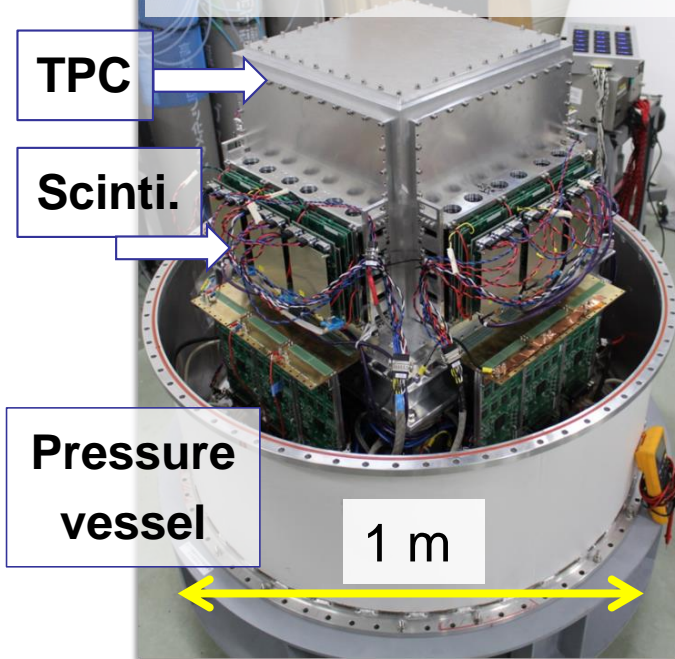
- Well-defined PSF without ML-EM
- Powerful BG rejection using  $dE/dx$ 
  - No shield => Wide field of view  $\sim 6\text{sr}$

T.Tanimori et al., *ApJ* (2015) accepted,  
[arXiv: 1507.03850](https://arxiv.org/abs/1507.03850) [astro-ph.IM]



# Current and near-future performance

SMILE-II Flight model



## SMILE-III

- ✓ 40 cm-cubic
- ✓ CF<sub>4</sub> gas @3atm
- ✓ 3 R.L. scintillator

## SMILE-II

## Point Spread Function

SMILE-II, III: ARM 5° ⇒ 8-15°  
SPD 30-100°

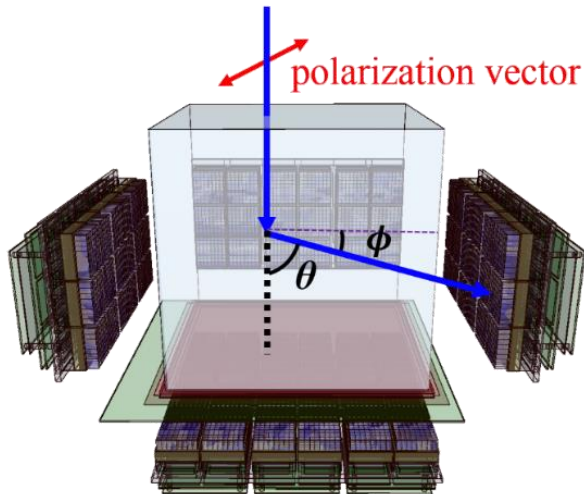
Near future: ARM 2° ⇒ ~1°  
SPD < 10°

Gas	Ar-based @1atm
Gas volume	30 cm-cubic
Track sampling	800 μm
Scintillator	pixel GSO:Ce (6 x 6 x 13 mm <sup>3</sup> ) 1 Radiation length
Energy resolution	10% @662 keV (FWHM)

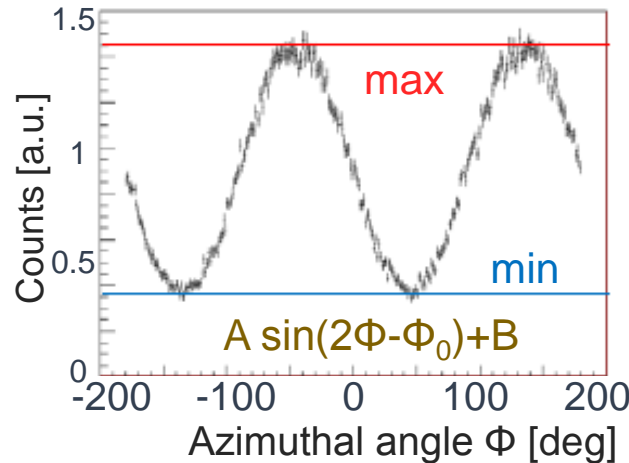
- SPDの改善策 25aSJ-11 宮本
- PSFの導入, 予想感度 25aSJ-12 高田

# as a Polarimeter

linearly polarized  $\gamma$ -ray



Asymmetric distribution of the scattered photons



$$\text{Modulation Factor} = \frac{\text{max} - \text{min}}{\text{max} + \text{min}}$$

~ 0.6 @ 200 keV,  
~ 0.5 @ 500 keV  
(simulation)

Minimum detectable polarization (MDP)

$$MDP_{[\%]} = \frac{429}{ASM} \sqrt{\frac{AS + B}{T}} \quad 99\% \text{ CL}$$

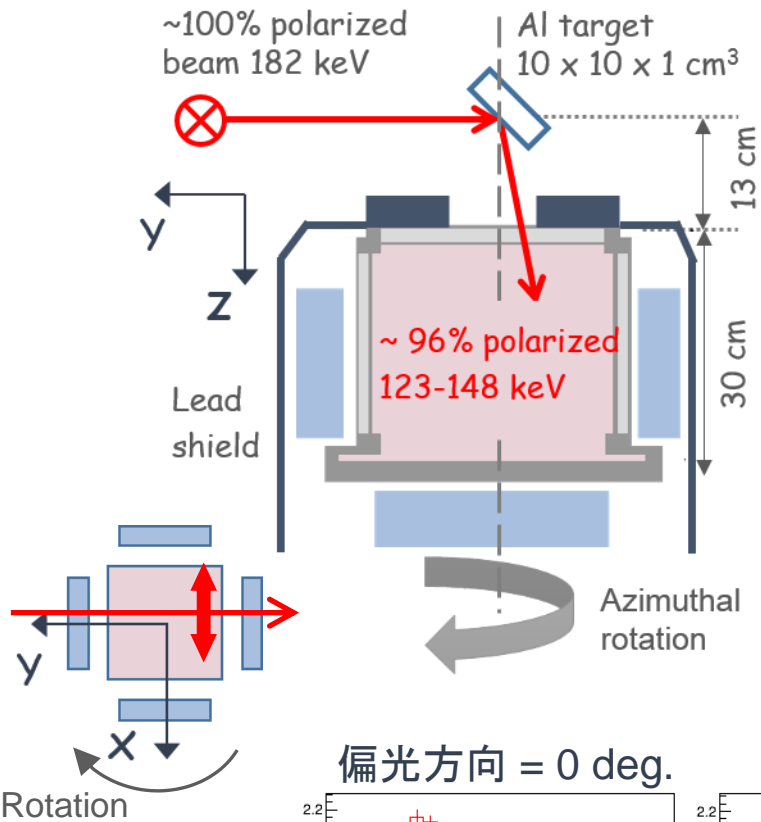
$$B \gg AS \Rightarrow MDP \propto \frac{\sqrt{B}}{AS}$$

A	Effective area [cm <sup>2</sup> ]	S	Signal [cm <sup>-2</sup> sec <sup>-1</sup> ]
M	Modulation Factor	B	Background [sec <sup>-1</sup> ]
T	Observation time [sec]		

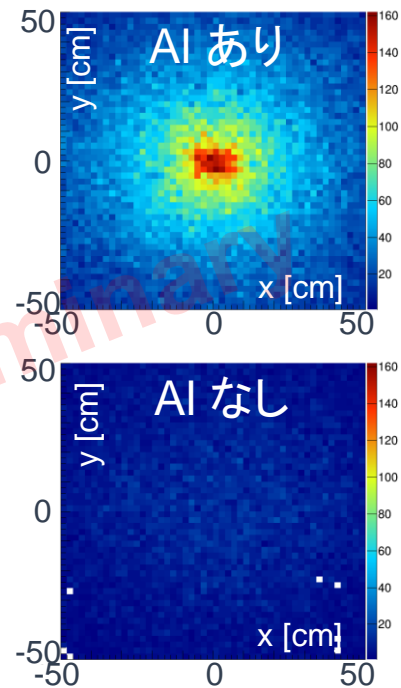
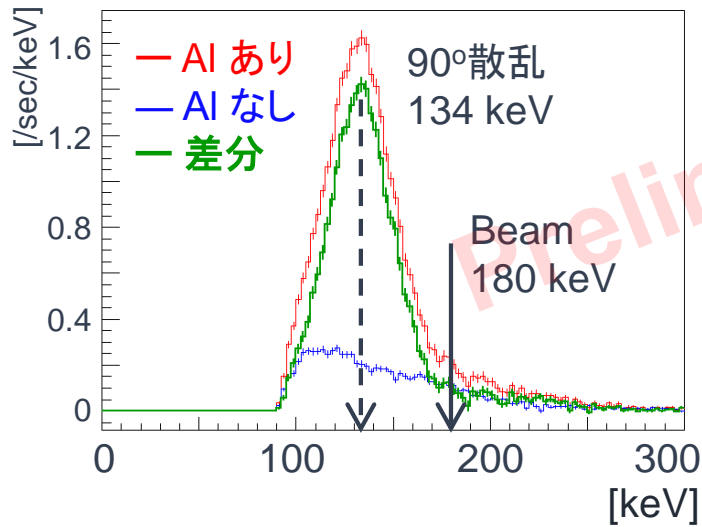
Sensitivity is limited by the background rate.

**Large advantage**  $\left\{ \begin{array}{l} \checkmark \text{ Powerful background rejection} \\ \checkmark \text{ Imaging with wide FOV } \sim 6\text{sr} \end{array} \right.$  faint sources, transient objects (GRBs)

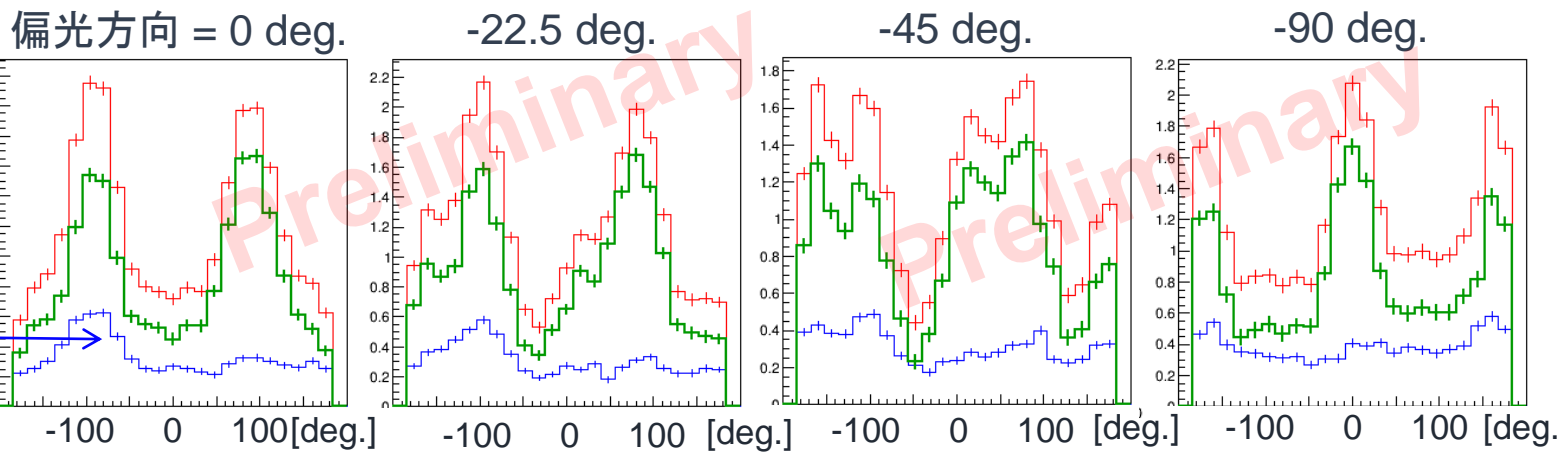
# Beam test@ SPring-8 BL08W



再構成スペクトル&イメージ (Al散乱事象)



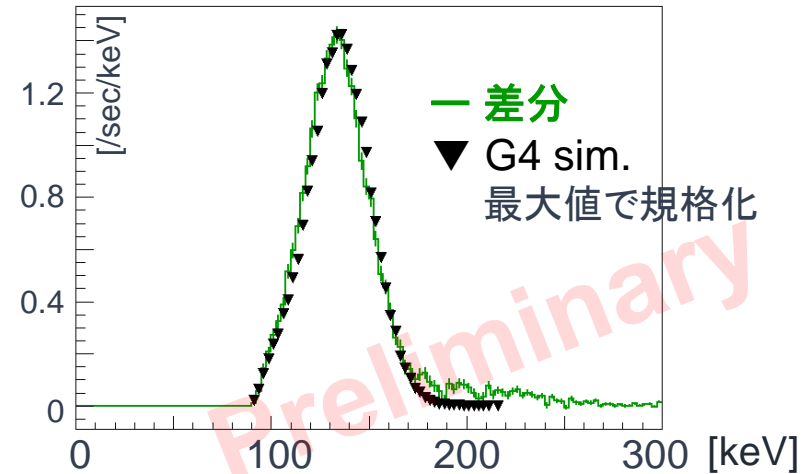
散乱ガンマ線方位角分布



# Geant4 Simulation

Geant4	Geant4 10.1 Patch-02
物理モデル	G4EmLivermorePolarizedPhysics (Fixed G4UrbanMscModel)
ジオメトリ	SMILE-II フライトモデル準拠
位置分解能	Scinti.のpixel sizeのみ
エネルギー分解能	TPC 22%@22 keV, Scinti. 10%@662 keV のガウス分布
その他	回路応答、解析特性は未考慮

再構成スペクトル  
(AI散乱事象)



偏光方向 = 0 deg.

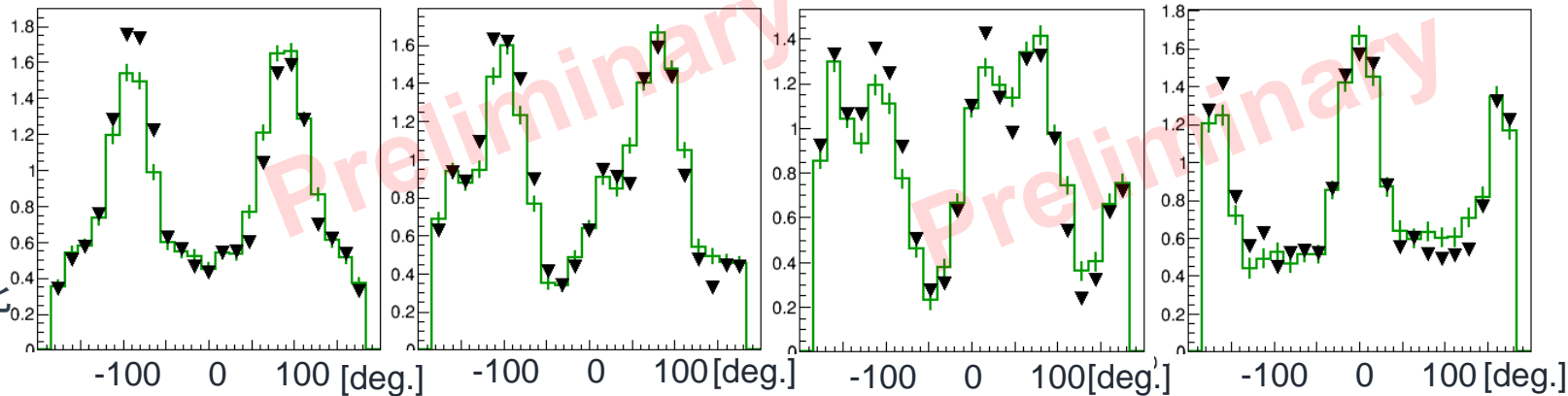
-22.5 deg.

-45 deg.

-90 deg.

散乱ガンマ線  
方位角分布

▼ G4 sim.  
面積で規格化





# Results

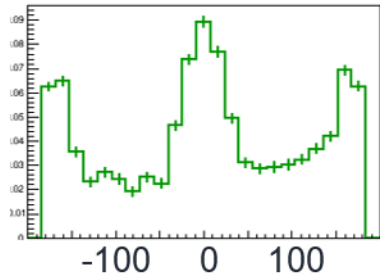
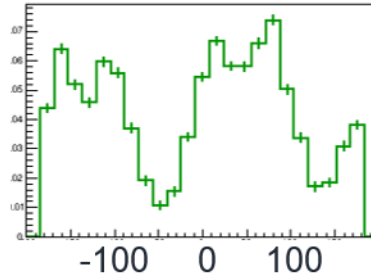
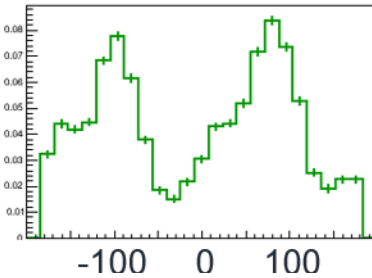
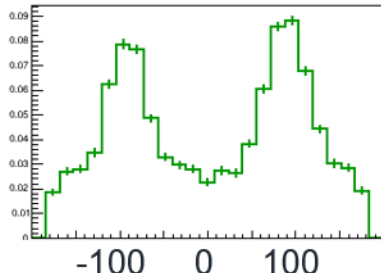
偏光方向 = 0 deg.

-22.5 deg.

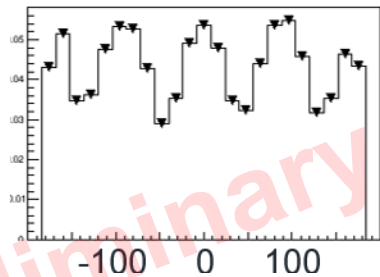
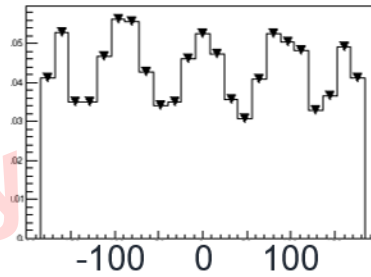
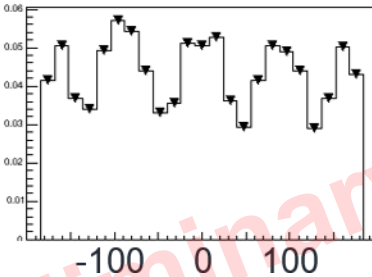
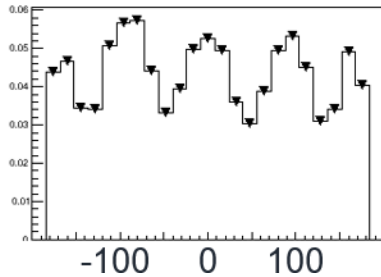
-45 deg.

-90 deg.

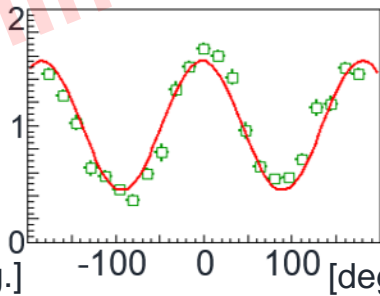
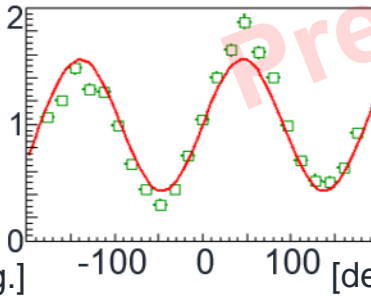
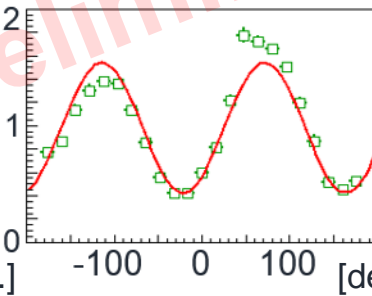
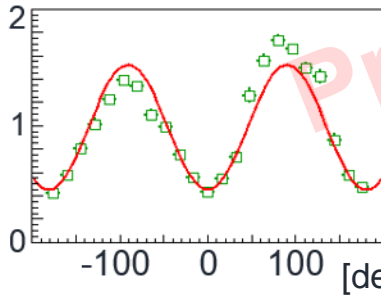
実験値  
(偏光)



Simulation  
無偏光  
レスポンス



幾何学補正  
実験値/Sim.



偏光方向 [deg.]  $0.19 \pm 1.03$

$-18.8 \pm 0.9$

$-44.4 \pm 0.7$

$-92.1 \pm 0.9$

Modulation Factor  $0.57 \pm 0.01$

$0.59 \pm 0.01$

$0.59 \pm 0.02$

$0.58 \pm 0.02$

◆ 偏光方向は設定値と誤差の範囲内で一致 ◆  $M \sim 0.6 @ 130 \text{ keV}$

# Summary

- ❑ Well-defined PSF without ML-EM
- ❑ Powerful background rejection by  $dE/dx$
- ❑ Polarization Measurement
  - $M \sim 0.6 @ 200 \text{ keV}$ ,  $\sim 0.5 @ 500 \text{ keV}$  (simulation)
  - Future Plan: SMILE-III ETCC ( $\sim 20 \text{ cm}^2 @ 200 \text{ keV}$ )
    - ✓ Polarization sensitivity :  $3\sigma$  MDP
    - Crab nebula  $\sim 15 \%$ , Cyg X-1  $\sim 20 \%$  (half-day flight)
    - GRBs  $\sim 6\%$  for  $10^{-6} \text{ erg/cm}^2 \text{ s}$  (2-3 GRBs/month)
    - $\sim 20\%$  for  $10^{-7} \text{ erg/cm}^2 \text{ s}$  ( $\sim 10$  GRBs/month)
- Beam test @ SPring-8
  - ✓ 撮像・分光・偏光の同時測定
  - ✓ シミュレーションによる幾何学補正
  - ✓ 検出偏光方向はセットアップと一致
  - ✓  $M \sim 0.6 @ 130 \text{ keV}$

