Performance Improvement of an Electron-Tracking Compton Camera by a New Track Reconstruction Method

S. Komura, T. Tanimori, H. Kubo, A. Takada, J. D. Parker, T. Mizumoto, Y. Mizumura, S. Sonoda, D. Tomono, S. Iwaki, T. Sawano, K. Nakamura, Y. Matsuoka, Y. Sato, S. Nakamura, M. Oda, K. Miuchi¹, S. Kabuki², Y. Kishimoto³, S. Kurosawa⁴, M. Tanaka³, M. Ikeno³, T. Uchida³

Kyoto University, ¹Kobe University, ²Tokai University, ³KEK, ⁴Tohoku University

1. Introduction



Targets in Sub-MeV/MeV gamma-ray band

- GRBs, nucleosynthesis in supernovae, particle acceleration in AGN, etc
- **Previous observations**
- COMPTEL (CGRO) Classical Compton Imaging
- > IBIS, SPI (INTEGRAL) Coded Aperture Imaging the problem of high background
- \Rightarrow low sensitivity

Background reduction is IMPORTANT

- **Our Detector** [1]
 - Electron Tracking Compton Imaging - Powerful background rejection ability

• Our Observations

- **SMILE** (Sub-MeV gamma ray Imaging Loaded-on-balloon Experiment)
- SMILE-I (2006) 10cm-cube ETCC @ Sanriku, Japan Observation of diffuse cosmic / atmospheric gamma rays Effective area ~ 0.01 cm² (150 keV-1.5 MeV) [2] Effective area must be
- SMILE-II (2014-) 30cm-cube ETCC improved ~ 50 times. Observation of Crab nebula / Cyg X-1 **Required effective area > 0.5 cm²** (@ angular resolution ~ 10°)

To obtain the required effective area for SMILE-II, we have improved the electron track reconstruction method by updating the track data encoding logic and developed a simple track analysis for the new logic. In this poster, we report the performance improvement of ETCCs by using the new track reconstruction method.

4. Performance improvement of µ-TPC



Schematic view of the ETCC

2. Our Detector : Electron Tracking Compton Camera Principle

- \succ Electron Tracker : μ -TPC Time Projection Chamber based on µ-PIC <u>**3D track**</u> and energy deposit of a recoil electron
- Photo Absorber : PSAs

Pixel Scintilater Arrays Energy and direction of a scattered gamma ray

- ✓ Reconstruction of incident photon event by event
- ✓ Kinematical background rejection ✓ Particle identification by dE/dX (BG rejection)
- \checkmark Large field of view (~3 str)

ψ-TPC



µ-PIC (Micro Pixel Chamber) [3] A gaseous 2D imaging detector with strip read out

 \checkmark Detection area : 10 x 10 cm², 30 x 30 cm² ✓ First amplifier : Gas Electron Multiplier (GEM) ✓ Typical gas gain : ~ 3000 (u-PIC) x 10 (GEM)

2D-readout (µ-PIC 400 µm pitch) + Drift time (100 MHz)



4 channels readout with resistor chain in each PMT

 \checkmark Hit pixel is determined by center of gravity of 4 ADC data.

GSO:Ce (6.71 g/cm³) Scintillator

• Experiment data of 10cm-cube μ -TPC (irradiate ¹³⁷Cs)



5. Simple track analysis for the new encoding logic

To reconstruct the incident gamma-ray, we need the information of the Compton scattered point and the recoil direction using the obtained 2D-track, (X,Z) and (Y,Z).

Reconstruction of 3D-track

Coincide with obtained 2D-track in an adequate gate (10-20 ns) using the software.

- **Determination of Compton scattered point** (2)Simple kinematic method. The closest hit point by XY-plane distance
- from the absorption point of scattered gamma ray.
- Determination of the recoil direction (3)



Pixel size	6 x 6 x 13 mm ³
Dynamic range	80 keV-1.3 MeV
Energy Resolution	10% @ 662 keV (FWHM)

3. Updating the track data encoding logic

All the signals from the anode (X) and cathode (Y) strips on the μ -PIC were discriminated, and they are sent to the following **FPGA encoding module** with an internal clock of 100 MHz.



The direction connecting the scattered point to the hit point which is farthest from the scattered point within 5 mm.



6. Performance improvement of an ETCC

Experiment Data of 10cm-cube ETCC (irradiate ¹³⁷Cs source at 60 cm above)

> Demonstration of Gamma-ray Reconstruction using a new logic



[3] T. Nagayoshi, Ph. D. Thesis, Kyoto Univ. (2004).

signals with 100 MHz clock.

ii. Coincide with detected anode(X) and cathode(Y) positions in one clock interval 10 nsec.

Record only the two edge positions 111. among the coincident hit positions with the hit timing.

Encoded Track Data $(X_{\min}, Y_{\min}, X_{\max}, Y_{\max}, T)$

A 10 nsec gate is too restricted for the coincidence of μ -PIC signals, and it caused the considerable loss of hit points (~70%) due to the time-walk of preamplifiers and the delay on FPGAs. This loss degraded the efficiency of an ETCC [6].

signals with 100 MHz clock.

Record all of the detected positions with the hit timing. Don't take coincidence.

Encoded Track Data (X, T), (Y, T)

The electron track are expected to be measured more precisely. In addition, we are able to estimated the energy deposition using the **Time Over Threshold (TOT)**, time between leading and trailing edges.

7. Summary & Future Work 10² Simulated Effective Area Measured efficiencies are 5-10 times higher [cm²] than using the previous logic and looked consistent with the simulated efficiency. D Measured angular resolutions, ARM and SPD, ArC2H6(90:10), 1 atm ArC2H6(90:10), 1.5 atm ArC2H6(90:10), 2.0 atm were improved remarkably. pureCF4, 2.0 atm pureCF4, 3.0 atm $(ARM : 8.2^{\circ} \rightarrow 5.8^{\circ}, SPD : 90^{\circ} - 200^{\circ} \rightarrow 93^{\circ} @ 662 \text{ keV})$ 10⁻² Energy [keV] ⇒ the 30cm-cube ETCC for SMILE-II will achieve 1cm² effective area and more than 3 times better sensitivity. ⇒ This sensitivity enables us to detect Crab nebula with 5 sigma level for several-hour observation [8]. □ In a few years, we will achieve 10 cm² effective area by changing the kind of gas and increasing the gas pressure. [4] H. Nishimura et al. NIM A 573 (2007) 115. 8. Reference [5] H. Kubo et al., IEEE NSS Conference Record (2005) 371. [1] T. Tanimori et al., New Astron. Rev., 48(2004), 263 [6] K. Hattori, Ph. D. Thesis, Kyoto Univ. (2009). [2] A. Takada et al., ApJ, 733 (2011), 13.

[7] K. Ueno, Ph. D. Thesis, Kyoto Univ. (2010). [8] T. Mizumoto, 2013 IEEE NSS Presentation (2013).

 10^{3}